# Package 'LSTS'

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R topics documented:
block.smooth.periodogram

**20** 

LS.summary																•
LS.whittle																
LS.whittle.loglik																
LS.whittle.loglik.sd .																
LS.whittle.loglik.theta																
malleco																
periodogram																
smooth.periodogram																
spectral.density																
ts.diag																

block.smooth.periodogram

Smooth Periodogram by Blocks

### Description

Plots the contour plot of the smoothing periodogram of a time series, by blocks or windows.

### Usage

**Index** 

```
block.smooth.periodogram(
   y,
   x = NULL,
   N = NULL,
   S = NULL,
   p = 0.25,
   spar.freq = 0,
   spar.time = 0
)
```

### Arguments

У	(type: numeric) data vector
X	(type: numeric) optional vector, if x = NULL then the function uses $(1,\ldots,n)$ where n is the length of y.
N	(type: numeric) value corresponding to the length of the window to compute periodogram. If N=NULL then the function will use $N={\rm trunc}(n^{0.8})$ , see Dahlhaus and Giraitis (1998) where $n$ is the length of the y vector.
S	(type: numeric) value corresponding to the lag with which will be taking the blocks or windows to calculate the periodogram.
p	(type: numeric) value used if it is desired that S is proportional to N. By default p=0.25, if S and N are not entered.
spar.freq	(type: numeric) smoothing parameter, typically (but not necessarily) in $(0,1]$ .
spar.time	(type: numeric) smoothing parameter, typically (but not necessarily) in $(0, 1]$ .

Box.Ljung.Test 3

#### **Details**

The number of windows of the function is  $m = \operatorname{trunc}((n-N)/S+1)$ , where trunc truncates de entered value and n is the length of the vector y. All windows are of the same length N, if this value isn't entered by user then is computed as  $N = \operatorname{trunc}(n^{0.8})$  (Dahlhaus). LSTS\_spb computes the periodogram in each of the M windows and then smoothes it two times with smooth.spline function; the first time using spar.freq parameter and the second time with spar.time. These windows overlap between them.

#### Value

A ggplot object.

#### References

For more information on theoretical foundations and estimation methods see Dahlhaus R, others (1997). "Fitting time series models to nonstationary processes." *The annals of Statistics*, **25**(1), 1–37. Dahlhaus R, Giraitis L (1998). "On the optimal segment length for parameter estimates for locally stationary time series." *Journal of Time Series Analysis*, **19**(6), 629–655.

#### See Also

```
arima.sim
```

#### **Examples**

block.smooth.periodogram(malleco)

Box.Ljung.Test

Ljung-Box Test Plot

#### **Description**

Plots the p-values Ljung-Box test.

#### Usage

```
Box.Ljung.Test(z, lag = NULL, main = NULL)
```

#### Arguments

z (type: numeric) data vector

lag (type: numeric) the number of periods for the autocorrelation

main (type: character) a title for the returned plot

4 hessian

#### **Details**

The Ljung-Box test is used to check if exists autocorrelation in a time series. The statistic is

$$q = n(n+2) \cdot \sum_{j=1}^{h} \hat{\rho}(j)^2 / (n-j)$$

with n the number of observations and  $\hat{\rho}(j)$  the autocorrelation coefficient in the sample when the lag is j. LSTS\_1btp computes q and returns the p-values graph with lag j.

#### Value

A ggplot object.

#### References

For more information on theoretical foundations and estimation methods see Brockwell PJ, Davis RA, Calder MV (2002). *Introduction to time series and forecasting*, volume 2. Springer. Ljung GM, Box GE (1978). "On a measure of lack of fit in time series models." *Biometrika*, **65**(2), 297–303.

#### See Also

periodogram

#### **Examples**

```
Box.Ljung.Test(malleco, lag = 5)
```

hessian

Hessian Matrix

### Description

Numerical approximation of the Hessian of a function.

#### **Usage**

```
hessian(f, x0, ...)
```

#### **Arguments**

f (type: numeric) name of function that defines log likelihood (or negative of it).

x0 (type: numeric) scalar or vector of parameters that give the point at which you

want the hessian estimated (usually will be the mle).

. . . Additional arguments to be passed to the function.

LS.kalman 5

#### **Details**

Computes the numerical approximation of the Hessian of f, evaluated at x0. Usually needs to pass additional parameters (e.g. data). N.B. this uses no numerical sophistication.

#### Value

An  $n \times n$  matrix of 2nd derivatives, where n is the length of x0.

#### See Also

```
arima.sim
```

#### **Examples**

```
# Variance of the maximum likelihood estimator for mu parameter in
# gaussian data
loglik <- function(series, x, sd = 1) {
    -sum(log(dnorm(series, mean = x, sd = sd)))
}
sqrt(c(var(malleco) / length(malleco), diag(solve(hessian(
    f = loglik, x = mean(malleco), series = malleco,
    sd = sd(malleco)
)))))</pre>
```

LS.kalman

Kalman filter for locally stationary processes

#### **Description**

This function run the state-space equations for expansion infinite of moving average in processes LS-ARMA or LS-ARFIMA.

#### Usage

```
LS.kalman(
    series,
    start,
    order = c(p = 0, q = 0),
    ar.order = NULL,
    ma.order = NULL,
    sd.order = NULL,
    d.order = NULL,
    include.d = FALSE,
    m = NULL
)
```

6 LS.kalman

### **Arguments**

series	(type: numeric) univariate time series.
start	(type: numeric) numeric vector, initial values for parameters to run the model.
order	(type: numeric) vector corresponding to ARMA model entered.
ar.order	(type: numeric) AR polimonial order.
ma.order	(type: numeric) MA polimonial order.
sd.order	(type: numeric) polinomial order noise scale factor.
d.order	(type: numeric) d polinomial order, where d is the ARFIMA parameter.
include.d	(type: numeric) logical argument for ARFIMA models. If include . d=FALSE then the model is an ARMA process.
m	(type: numeric) truncation order of the MA infinity process. By default $m=0.25n^{0.8}$ where n the length of series.

#### **Details**

The model fit is done using the Whittle likelihood, while the generation of innovations is through Kalman Filter. Details about ar.order, ma.order, sd.order and d.order can be viewed in LS.whittle.

### Value

#### A list with:

residuals standard residuals.

fitted\_values model fitted values.

delta variance prediction error.

#### References

For more information on theoretical foundations and estimation methods see Brockwell PJ, Davis RA, Calder MV (2002). *Introduction to time series and forecasting*, volume 2. Springer. Palma W (2007). *Long-memory time series: theory and methods*, volume 662. John Wiley & Sons. Palma W, Olea R, Ferreira G (2013). "Estimation and forecasting of locally stationary processes." *Journal of Forecasting*, **32**(1), 86–96.

#### **Examples**

```
fit_kalman <- LS.kalman(malleco, start(malleco))</pre>
```

LS.summary 7

LS.summary

Summary for Locally Stationary Time Series

### Description

Produces a summary of the results to Whittle estimator to Locally Stationary Time Series (LS. whittle function).

### Usage

```
LS.summary(object)
```

#### **Arguments**

object

(type: list) the output of LS. whittle function

#### **Details**

Calls the output from LS.whittle and computes the standard error and p-values to provide a detailed summary.

#### Value

A list with the following components:

summary a resume table with estimate, std. error, z-value and p-value of the model.

aic AIC of the model.

npar number of parameters in the model.

#### See Also

```
LS.whittle
```

### **Examples**

```
fit_whittle <- LS.whittle(
  series = malleco, start = c(1, 1, 1, 1),
  order = c(p = 1, q = 0), ar.order = 1, sd.order = 1, N = 180, n.ahead = 10
)
LS.summary(fit_whittle)</pre>
```

LS.whittle

LS.whittle

Whittle estimator to Locally Stationary Time Series

### Description

This function computes Whittle estimator to LS-ARMA and LS-ARFIMA models.

### Usage

```
LS.whittle(
  series,
  start,
  order = c(p = 0, q = 0),
  ar.order = NULL,
  ma.order = NULL,
  sd.order = NULL,
  d.order = NULL,
  include.d = FALSE,
  N = NULL
  S = NULL,
  include.taper = TRUE,
  control = list(),
  lower = -Inf,
  upper = Inf,
  m = NULL,
  n.ahead = 0
)
```

#### **Arguments**

series	(type: numeric) univariate time series.
start	(type: numeric) numeric vector, initial values for parameters to run the model.
order	(type: numeric) vector corresponding to ARMA model entered.
ar.order	(type: numeric) AR polimonial order.
ma.order	(type: numeric) MA polimonial order.
sd.order	(type: numeric) polinomial order noise scale factor.
d.order	(type: numeric) d polinomial order, where d is the ARFIMA parameter.
include.d	(type: numeric) logical argument for ARFIMA models. If include $.d=FALSE$ then the model is an ARMA process.
N	(type: numeric) value corresponding to the length of the window to compute periodogram. If N=NULL then the function will use $N={\rm trunc}(n^{0.8})$ , see Dahlhaus (1998) where $n$ is the length of the y vector.
S	(type: numeric) value corresponding to the lag with which will go taking the blocks or windows.

LS.whittle

include.taper (type: logical) logical argument that by default is TRUE. See periodogram.

control (type: list) A list of control parameters. More details in nlminb.

lower (type: numeric) lower bound, replicated to be as long as start. If unspecified,

all parameters are assumed to be lower unconstrained.

upper (type: numeric) upper bound, replicated to be as long as start. If unspecified,

all parameters are assumed to be upper unconstrained.

m (type: numeric) truncation order of the MA infinity process, by default m =

 $0.25n^{0.8}$ . Parameter used in LSTS\_kalman.

n.ahead (type: numeric) The number of steps ahead for which prediction is required. By

default is zero.

#### **Details**

This function estimates the parameters in models: LS-ARMA

$$\Phi(t/T, B) Y_{t,T} = \Theta(t/T, B) \sigma(t/T) \varepsilon_t$$

and LS-ARFIMA

$$\Phi(t/T, B) Y_{t,T} = \Theta(t/T, B) (1 - B)^{-d(t/T)} \sigma(t/T) \varepsilon_t,$$

with infinite moving average expansion

$$Y_{t,T} = \sigma(t/T) \sum_{j=0}^{\infty} \psi(t/T) \, \varepsilon_t,$$

for  $t=1,\ldots,T$ , where for  $u=t/T\in[0,1]$ ,  $\Phi(u,B)=1+\phi_1(u)B+\cdots+\phi_p(u)B^p$  is an autoregressive polynomial,  $\Theta(u,B)=1+\theta_1(u)B+\cdots+\theta_q(u)B^q$  is a moving average polynomial, d(u) is a long-memory parameter,  $\sigma(u)$  is a noise scale factor and  $\{\varepsilon_t\}$  is a Gaussian white noise sequence with zero mean and unit variance. This class of models extends the well-known ARMA and ARFIMA process, which is obtained when the components  $\Phi(u,B)$ ,  $\Theta(u,B)$ , d(u) and  $\sigma(u)$  do not depend on u. The evolution of these models can be specified in terms of a general class of functions. For example, let  $\{g_j(u)\}, j=1,2,\ldots$ , be a basis for a space of smoothly varying functions and let  $d_{\theta}(u)$  be the time-varying long-memory parameter in model LS-ARFIMA. Then we could write  $d_{\theta}(u)$  in terms of the basis  $\{g_j(u)=u^j\}$  as follows  $d_{\theta}(u)=\sum_{j=0}^k \alpha_j\,g_j(u)$  for unknown values of k and  $\theta=(\alpha_0,\alpha_1,\ldots,\alpha_k)'$ . In this situation, estimating  $\theta$  involves determining k and estimating the coefficients  $\alpha_0,\alpha_1,\ldots,\alpha_k$ . LS. whittle optimizes LS. whittle.loglik as objective function using nlminb function, for both LS-ARMA (include.d=FALSE) and LS-ARFIMA (include.d=TRUE) models. Also computes Kalman filter with LS. kalman and this values are given in var.coef in the output.

#### Value

A list with the following components:

coef The best set of parameters found.

var.coef covariance matrix approximated for maximum likelihood estimator  $\hat{\theta}$  of  $\theta := (\theta_1, \dots, \theta_k)'$ . This matrix is approximated by  $H^{-1}/n$ , where H is the Hessian

matrix  $[\partial^2 \ell(\theta)/\partial \theta_i \partial \theta_j]_{i,j=1}^k$ .

LS.whittle

log-likelihood of coef, calculated with LS. whittle. loglik aic Akaike'S 'An Information Criterion', for one fitted model LS-ARMA or LS-ARFIMA. The formula is -2L + 2k/n, where L represents the log-likelihood, k represents the number of parameters in the fitted model and n is equal to the length of the series. original time serie. series residuals standard residuals. model fitted values. fitted.values predictions of the model. pred the estimated standard errors. se A list representing the fitted model. model

#### See Also

nlminb, LS.kalman

#### **Examples**

```
# Analysis by blocks of phi and sigma parameters
N <- 200
S <- 100
M <- trunc((length(malleco) - N) / S + 1)
table <- c()
for (j in 1:M) {
  x \leftarrow malleco[(1 + S * (j - 1)):(N + S * (j - 1))]
  table <- rbind(table, nlminb(</pre>
    start = c(0.65, 0.15), N = N,
    objective = LS.whittle.loglik,
    series = x, order = c(p = 1, q = 0)
  )$par)
}
u \leftarrow (N / 2 + S * (1:M - 1)) / length(malleco)
table <- as.data.frame(cbind(u, table))</pre>
colnames(table) <- c("u", "phi", "sigma")</pre>
# Start parameters
phi <- smooth.spline(table$phi, spar = 1, tol = 0.01)$y</pre>
fit.1 <- nls(phi \sim a0 + a1 * u, start = list(a0 = 0.65, a1 = 0.00))
sigma <- smooth.spline(table$sigma, spar = 1)$y</pre>
fit.2 <- nls(sigma \sim b0 + b1 * u, start = list(b0 = 0.65, b1 = 0.00))
fit_whittle <- LS.whittle(</pre>
  series = malleco, start = c(coef(fit.1), coef(fit.2)), order = c(p = 1, q = 0),
  ar.order = 1, sd.order = 1, N = 180, n.ahead = 10
```

LS.whittle.loglik

LS.whittle.loglik

Locally Stationary Whittle log-likelihood Function

### Description

This function computes Whittle estimator for LS-ARMA and LS-ARFIMA models, in data with mean zero. If mean is not zero, then it is subtracted to data.

### Usage

```
LS.whittle.loglik(
    x,
    series,
    order = c(p = 0, q = 0),
    ar.order = NULL,
    ma.order = NULL,
    sd.order = NULL,
    d.order = NULL,
    include.d = FALSE,
    N = NULL,
    S = NULL,
    include.taper = TRUE
)
```

#### **Arguments**

x	(type: numeric) parameter vector.
series	(type: numeric) univariate time series.
order	(type: numeric) vector corresponding to ARMA model entered.
ar.order	(type: numeric) AR polimonial order.
ma.order	(type: numeric) MA polimonial order.
sd.order	(type: numeric) polinomial order noise scale factor.
d.order	(type: numeric) d polinomial order, where d is the ARFIMA parameter.
include.d	(type: numeric) logical argument for ARFIMA models. If include.d=FALSE then the model is an ARMA process.
N	(type: numeric) value corresponding to the length of the window to compute periodogram. If N=NULL then the function will use $N={\rm trunc}(n^{0.8})$ , see Dahlhaus (1998) where $n$ is the length of the y vector.
S	(type: numeric) value corresponding to the lag with which will go taking the blocks or windows.
include.taper	(type: logical) logical argument that by default is TRUE. See periodogram.

12 LS.whittle.loglik.sd

#### **Details**

The estimation of the time-varying parameters can be carried out by means of the Whittle log-likelihood function proposed by Dahlhaus (1997),

$$L_n(\theta) = \frac{1}{4\pi} \frac{1}{M} \int_{-\pi}^{\pi} \left\{ log f_{\theta}(u_j, \lambda) + \frac{I_N(u_j, \lambda)}{f_{\theta}(u_j, \lambda)} \right\} d\lambda$$

where M is the number of blocks, N the length of the series per block, n = S(M-1) + N, S is the shift from block to block,  $u_j = t_j/n$ ,  $t_j = S(j-1) + N/2$ ,  $j = 1, \ldots, M$  and  $\lambda$  the Fourier frequencies in the block  $(2\pi k/N, k = 1, \ldots, N)$ .

#### References

For more information on theoretical foundations and estimation methods see Brockwell PJ, Davis RA, Calder MV (2002). *Introduction to time series and forecasting*, volume 2. Springer. Palma W, Olea R, others (2010). "An efficient estimator for locally stationary Gaussian long-memory processes." *The Annals of Statistics*, **38**(5), 2958–2997.

#### See Also

nlminb, LS.kalman

LS.whittle.loglik.sd Locally Stationary Whittle Log-likelihood sigma

#### Description

This function calculates log-likelihood with known  $\theta$ , through LS.whittle.loglik function.

#### Usage

```
LS.whittle.loglik.sd(
    x,
    series,
    order = c(p = 0, q = 0),
    ar.order = NULL,
    ma.order = NULL,
    sd.order = NULL,
    d.order = NULL,
    include.d = FALSE,
    N = NULL,
    include.taper = TRUE,
    theta.par = numeric()
)
```

LS.whittle.loglik.theta

#### **Arguments**

X	(type: numeric) parameter vector.
series	(type: numeric) univariate time series.
order	(type: numeric) vector corresponding to ARMA model entered.
ar.order	(type: numeric) AR polimonial order.
ma.order	(type: numeric) MA polimonial order.
sd.order	(type: numeric) polinomial order noise scale factor.
d.order	(type: numeric) d polinomial order, where d is the ARFIMA parameter.
include.d	(type: numeric) logical argument for ARFIMA models. If include . d=FALSE then the model is an ARMA process.
N	(type: numeric) value corresponding to the length of the window to compute periodogram. If N=NULL then the function will use $N=\mathrm{trunc}(n^{0.8})$ , see Dahlhaus (1998) where $n$ is the length of the y vector.
S	(type: numeric) value corresponding to the lag with which will go taking the blocks or windows.
include.taper	(type: logical) logical argument that by default is TRUE. See periodogram.
theta.par	(type: numeric) vector with the known parameters of the model.

#### **Details**

This function computes LS.whittle.loglik with x as x = c(theta.par, x).

```
LS.whittle.loglik.theta
```

Locally Stationary Whittle Log-likelihood theta

### Description

Calculate the log-likelihood with  $\sigma$  known, through LS.whittle.loglik function.

### Usage

```
LS.whittle.loglik.theta(
    x,
    series,
    order = c(p = 0, q = 0),
    ar.order = NULL,
    ma.order = NULL,
    sd.order = NULL,
    d.order = NULL,
    include.d = FALSE,
    N = NULL,
    include.taper = TRUE,
    sd.par = 1
)
```

14 malleco

#### **Arguments**

X	(type: numeric) parameter vector.
series	(type: numeric) univariate time series.
order	(type: numeric) vector corresponding to ARMA model entered.
ar.order	(type: numeric) AR polimonial order.
ma.order	(type: numeric) MA polimonial order.
sd.order	(type: numeric) polinomial order noise scale factor.
d.order	(type: numeric) d polinomial order, where d is the ARFIMA parameter.
include.d	(type: numeric) logical argument for ARFIMA models. If include . d=FALSE then the model is an ARMA process.
N	(type: numeric) value corresponding to the length of the window to compute periodogram. If N=NULL then the function will use $N=\mathrm{trunc}(n^{0.8})$ , see Dahlhaus (1998) where $n$ is the length of the y vector.
S	(type: numeric) value corresponding to the lag with which will go taking the blocks or windows.
include.taper	(type: logical) logical argument that by default is TRUE. See periodogram.
sd.par	(type: numeric) value corresponding to known variance.

#### **Details**

This function computes LS.whittle.loglik with x as x = c(x, sd.par).

### Description

A ts object containing average annual ring width measured in milimiters for different Araucaria Araucana trees in the Malleco Region (Chile). The years of observation in this data cover the period 1242-1975.

#### **Format**

A time series object with 734 elements

### Author(s)

National Oceanic and Atmospheric Administration (NOAA)

periodogram 15

periodogram	Periodogram function	

#### **Description**

This function computes the periodogram from a stationary time serie. Returns the periodogram, its graph and the Fourier frequency.

#### Usage

```
periodogram(y, plot = TRUE, include.taper = FALSE)
```

#### **Arguments**

y (type: numeric) data vector

plot (type: logical) logical argument which allows to plot the periodogram. Defaults

to TRUE.

include.taper (type: logical) logical argument which by default is FALSE. If include.taper=TRUE

then y is multiplied by  $0.5(1 - \cos(2\pi(n-1)/n))$  (cosine bell).

#### **Details**

The tapered periodogram it is given by

$$I(\lambda) = \frac{|D_n(\lambda)|^2}{2\pi H_{2,n}(0)}$$

with  $D(\lambda) = \sum_{s=0}^{n-1} h\left(\frac{s}{N}\right) y_{s+1} e^{-i\lambda s}$ ,  $H_{k,n} = \sum_{s=0}^{n-1} h\left(\frac{s}{N}\right)^k e^{-i\lambda s}$  and  $\lambda$  are Fourier frequencies defined as  $2\pi k/n$ , with  $k=1,\ldots,n$ . The data taper used is the cosine bell function,  $h(x) = \frac{1}{2}[1-\cos(2\pi x)]$ . If the series has missing data, these are replaced by the average of the data and n it is corrected by \$n-N\$, where N is the amount of missing values of serie. The plot of the periodogram is periodogram values vs.  $\lambda$ .

#### Value

A list with with the periodogram and the lambda values.

#### References

For more information on theoretical foundations and estimation methods see Brockwell PJ, Davis RA, Calder MV (2002). *Introduction to time series and forecasting*, volume 2. Springer. Dahlhaus R, others (1997). "Fitting time series models to nonstationary processes." *The annals of Statistics*, **25**(1), 1–37.

#### See Also

fft, Mod, smooth.spline.

16 smooth.periodogram

#### **Examples**

```
# AR(1) simulated
set.seed(1776)
ts.sim <- arima.sim(n = 1000, model = list(order = c(1, 0, 0), ar = 0.7))
per <- periodogram(ts.sim)
per$plot</pre>
```

smooth.periodogram

Smoothing periodogram

#### **Description**

This function returns the smoothing periodogram of a stationary time serie, its plot and its Fourier frequency.

#### Usage

```
smooth.periodogram(y, plot = TRUE, spar = 0)
```

#### **Arguments**

y (type: numeric) data vector.

plot (type: logical) logical argument which allows to plot the periodogram. Defaults

to TRUE.

spar (type: numeric) smoothing parameter, typically (but not necessarily) in (0,1].

#### **Details**

smooth.periodogram computes the periodogram from y vector and then smooth it with *smoothing spline* method, which basically approximates a curve using a cubic spline (see more details in smooth.spline).  $\lambda$  is the Fourier frequency obtained through periodogram. It must have caution with the minimum length of y, because smooth.spline requires the entered vector has at least length 4 and the length of y does not equal to the length of the data of the periodogram that smooth.spline receives. If it presents problems with tol (**tol**erance), see smooth.spline.

#### Value

A list with with the smooth periodogram and the lambda values

#### See Also

```
smooth.spline,periodogram
```

spectral.density 17

#### **Examples**

```
# AR(1) simulated
require(ggplot2)
set.seed(1776)
ts.sim <- arima.sim(n = 1000, model = list(order = c(1, 0, 0), ar = 0.7))
per <- periodogram(ts.sim)
aux <- smooth.periodogram(ts.sim, plot = FALSE, spar = .7)
sm_p <- data.frame(x = aux$lambda, y = aux$smooth.periodogram)
sp_d <- data.frame(
    x = aux$lambda,
    y = spectral.density(ar = 0.7, lambda = aux$lambda)
)
g <- per$plot
g +
    geom_line(data = sm_p, aes(x, y), color = "#ff7f0e") +
    geom_line(data = sp_d, aes(x, y), color = "#d31244")</pre>
```

spectral.density

Spectral Density

#### **Description**

Returns theoretical spectral density evaluated in ARMA and ARFIMA processes.

#### Usage

```
spectral.density(ar = numeric(), ma = numeric(), d = 0, sd = 1, lambda = NULL)
```

#### **Arguments**

ar	(type: numeric) AR vector. If the time serie doesn't have AR term then omit it. For more details see the examples.
ma	(type: numeric) MA vector. If the time serie doesn't have MA term then omit it. For more details see the examples.
d	(type: numeric) Long-memory parameter. If d is zero, then the process is $ARMA(p,q)$ .
sd	(type: numeric) Noise scale factor, by default is 1.
lambda	(type: numeric) $\lambda$ parameter on which the spectral density is calculated/computed. If lambda=NULL then it is considered a sequence between 0 and $\pi$ .

#### **Details**

The spectral density of an ARFIMA(p,d,q) processes is

$$f(\lambda) = \frac{\sigma^2}{2\pi} \cdot \left(2\sin(\lambda/2)\right)^{-2d} \cdot \frac{\left|\theta\left(\exp\left(-i\lambda\right)\right)\right|^2}{\left|\phi\left(\exp\left(-i\lambda\right)\right)\right|^2}$$

18 ts.diag

With  $-\pi \le \lambda \le \pi$  and -1 < d < 1/2. |x| is the Mod of x. LSTS\_sd returns the values corresponding to  $f(\lambda)$ . When d is zero, the spectral density corresponds to an ARMA(p,q).

#### Value

An unnamed vector of numeric class.

#### References

For more information on theoretical foundations and estimation methods see Brockwell PJ, Davis RA, Calder MV (2002). *Introduction to time series and forecasting*, volume 2. Springer. Palma W (2007). *Long-memory time series: theory and methods*, volume 662. John Wiley & Sons.

#### **Examples**

```
# Spectral Density AR(1)
require(ggplot2)
f <- spectral.density(ar = 0.5, lambda = malleco)
ggplot(data.frame(x = malleco, y = f)) +
  geom_line(aes(x = as.numeric(x), y = as.numeric(y))) +
  labs(x = "Frequency", y = "Spectral Density") +
  theme_minimal()</pre>
```

ts.diag

Diagnostic Plots for Time Series fits

#### Description

Plot time-series diagnostics.

### Usage

```
ts.diag(x, lag = 10, band = qnorm(0.975)/sqrt(length(x)))
```

### **Arguments**

x (type: numeric) residuals of the fitted time series model.

lag (type: numeric) maximum lag at which to calculate the acf and Ljung-Box test.

By default set to 10.

band (type: numeric) absolute value for bandwidth in the the ACF plot. By default

set to 'qnorm(0.975)/sqrt(n)' which approximates to 0.07 for malleco data (n =

734)

#### Details

This function plot the residuals, the autocorrelation function of the residuals (ACF) and the p-values of the Ljung-Box Test for all lags up to lag.

ts.diag

### Value

A ggplot object.

### See Also

Box.Ljung.Test

### Examples

ts.diag(malleco)

## **Index**

```
* data
    malleco, 14
\operatorname{arima.sim}, 3, 5
\verb|block.smooth.periodogram|, 2
Box.Ljung.Test, 3, 19
fft, 15
hessian, 4
LS.kalman, 5, 9, 10, 12
LS.summary, 7
LS.whittle, 6, 7, 8, 10
LS.whittle.loglik, 9, 11, 13, 14
LS.whittle.loglik.sd, 12
LS.whittle.loglik.theta, 13
malleco, 14
Mod, 15, 18
nlminb, 9, 10, 12
periodogram, 4, 9, 11, 13, 14, 15, 16
smooth.periodogram, 16
smooth.spline, 3, 15, 16
spectral.density, 17
trunc, 3
ts.diag, 18
```